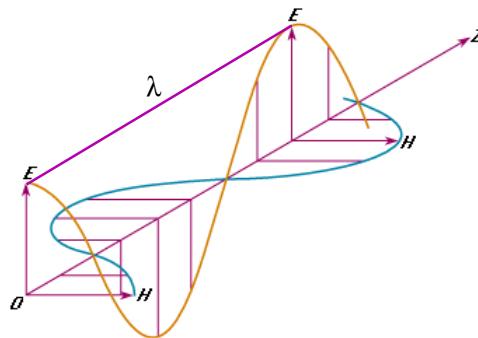


Characteristics of “free” em. Waves

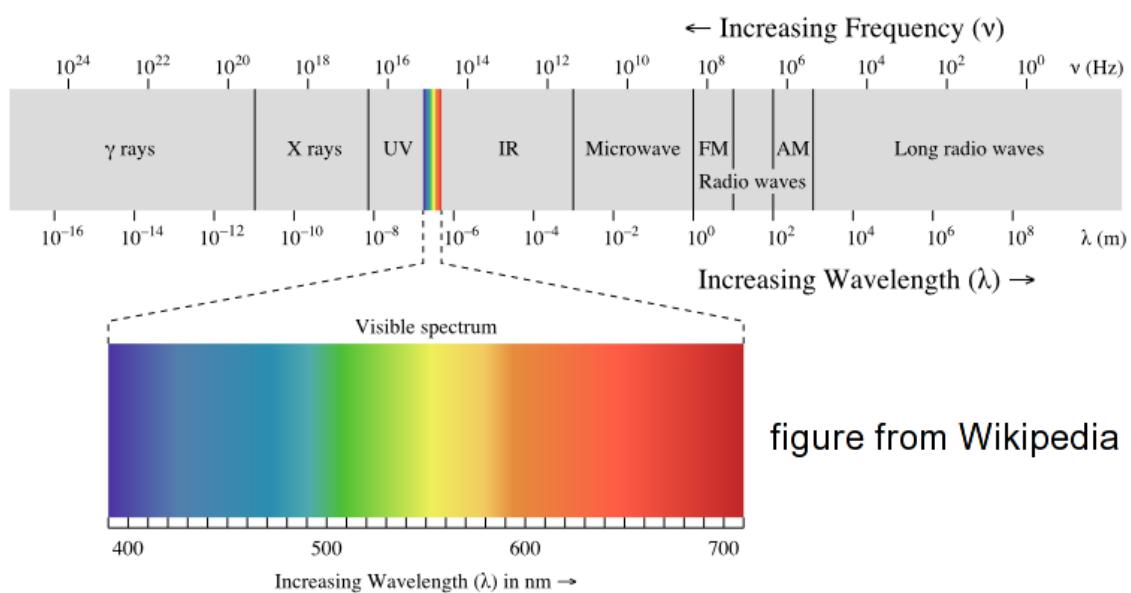
- ▶ electric and magnetic fields oscillate \perp to direction of propagation \Rightarrow **transverse waves**
- ▶ $\vec{E} \perp \vec{B}$ are in phase
- ▶ phase velocity: $c = 1/\sqrt{\mu_0\epsilon_0}$ = **speed of light**
- ▶ dispersion relation: $\omega = 2\pi f = c|\vec{k}|$ or $\lambda = cT$ ($|\vec{k}| = 2\pi/\lambda$)
- ▶ sources (not explained in this course)
 - ▶ accelerated charged particles =
 - ▶ **time-dependent** electric charges and currents
 - ▶ modern picture: quantum-mechanical transitions in atoms (visible light, UV) and nuclei (γ rays)



Hendrik van Hees

Electromagnetic waves

The em. spectrum



The em. spectrum

	f (Hz)	λ (m)	source (ex.)
γ rays	$> 10^{20}$	$< 10^{-12}$	radioactivity, nuclear transitions
X rays	$> 3 \cdot 10^{16}$	$< 10^{-8}$	bremsstrahlung radiation, atomic transitions
UV	$7.5 \cdot 10^{14} - 3 \cdot 10^{16}$	$4 \cdot 10^{-7} - 10^{-8}$	our Sun
visible light	$4 \cdot 10^{14} - 7.5 \cdot 10^{14}$	$4 \cdot 10^{-7} - 7.5 \cdot 10^{-7}$	atomic transitions
infrared	$3 \cdot 10^{11} - 4 \cdot 10^{14}$	$10^{-3} - 7.5 \cdot 10^{-7}$	transitions between vibrational modes of molecules

The em. spectrum

	f (Hz)	λ (m)	source (ex.)
Millimeter Waves	$30-300 \cdot 10^9$	$10^{-3}-10^{-2}$	antenna
microwaves	$1.6-30 \cdot 10^9$	$10-187 \cdot 10^{-3}$	magnetrons (microwave oven), rotation and torsion transitions of molecules, cosmic microwave background
radio waves	$5 \cdot 10^5 - 1.6 \cdot 10^9$	0.2-200	antenna