

Strangeness production in heavy-ion collisions at $E_{\text{kin}} = 1-2 \text{ AGeV}$

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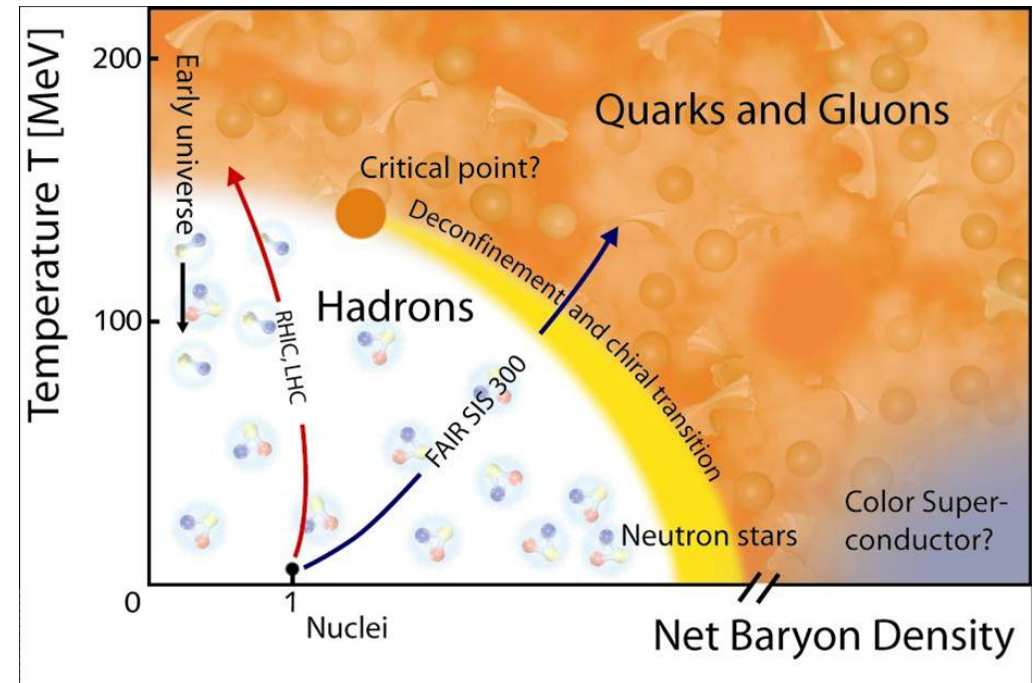
In collaboration with Hannah Petersen and Vinzent Steinberg



Motivation

- Nuclear matter under extreme conditions
 - State of the universe microseconds after the Big Bang
 - Hot and dense matter
 - Quark-Gluon-Plasma (QGP)
 - Quarks and gluons are **deconfined**
 - At high temperatures and densities hadronic matter forms QGP

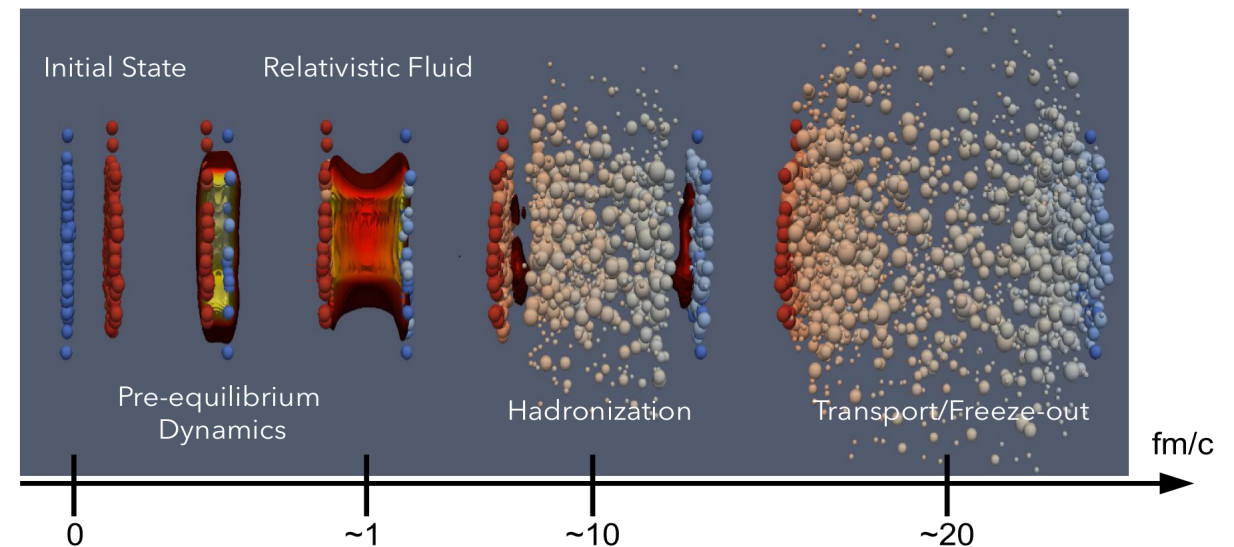
Goal: Understand phase transition



Source: teilchen.at

Motivation

- Heavy-ion collisions
 - At low energies, unique possibility to study nuclear matter at high densities
 - Created particles give information about conditions inside reaction zone
 - Sensitive probe → **strange particles**
- Specific production mechanism and long mean free path
 - ideal probes for exploring high density phase in heavy-ion reactions



Strangeness production

- Strange particles contain an s- or \bar{s} -quark
- Strange quarks have higher masses (95 MeV) than up and down quarks (2-5 MeV)
- Created by strong interaction and decay via weak interaction
→ long lifetime $1,238 \cdot 10^{-8}$ s
- Examples: K^+ ($u\bar{s}$), K^- ($\bar{u}s$), K^0 ($d\bar{s}$), \bar{K}^0 ($\bar{d}s$), Λ^0 (uds), Σ (uus)
- Production channels at low energies:
$$NN \rightarrow N\Lambda K^+$$
$$\pi Y \leftrightarrow K^- N \quad Y = \Lambda, \Sigma$$
- **Strangeness exchange** suggested as an important production channel in nucleus – nucleus collisions
 - Responsible for appearance and disappearance of K^-

Transport models

- Offer deeper understanding of the dynamics of heavy-ion collisions
- Simulate whole reaction
- Based on microscopic theory (transport theory)
- Whole information of particles is available in phase space
- Effectively solve relativistic Boltzmann equation:

$$p^\mu \partial_\mu f_i(x, p) + m_i F^\alpha \partial_\alpha^p f_i(x, p) = C_{coll}^i$$

SMASH

- Simulating **M**any **A**ccelerated **S**trongly-interacting **H**adrons
- C++ based hadronic transport code
- Simulation of nucleus-nucleus collisions at low energies
- Geometric collision criterion:

$$d_{trans} < d_{int} = \sqrt{\frac{\sigma_{tot}}{\pi}}$$

Strangeness production in SMASH

- Strange particles: K, K^* (11), Λ, Λ^* (13), Σ, Σ^* (9), Ξ, Ξ^* (5), Ω, Ω^* (1)
- Production via:

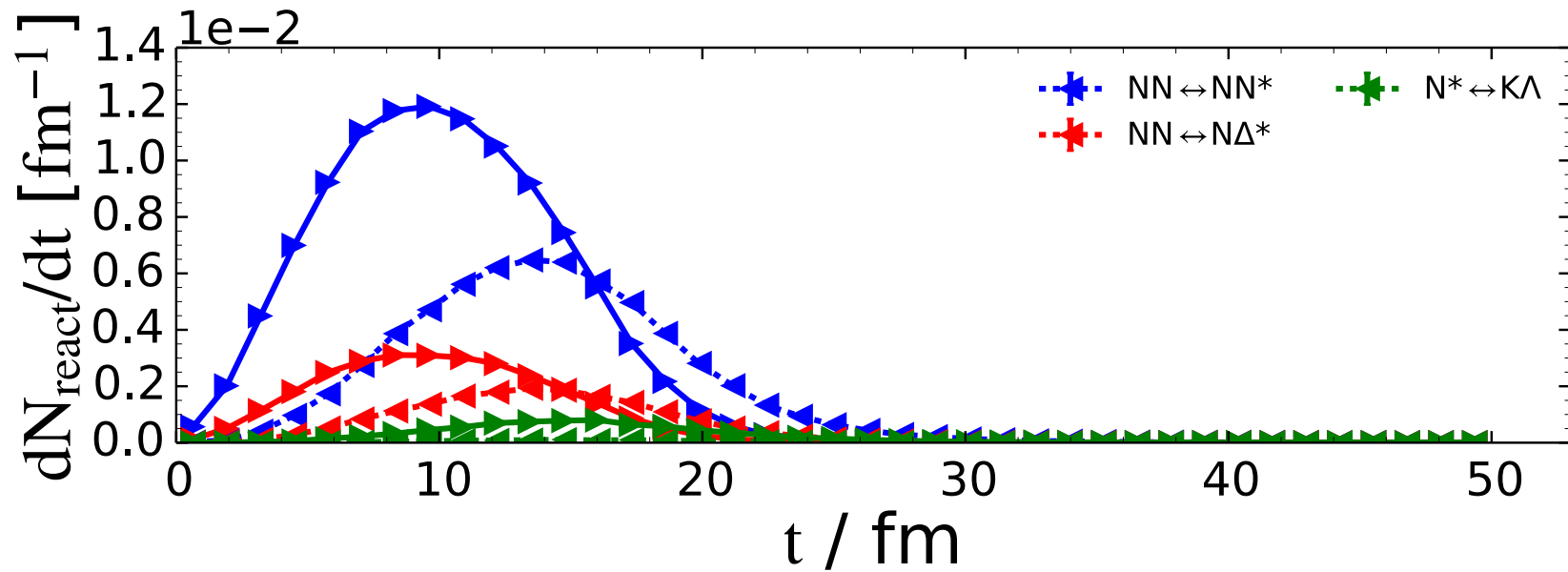
$$(1) NN \rightarrow NN^*, N\Delta^*$$

$$(2) N^*, \Delta^* \rightarrow YK \quad Y = \Lambda, \Sigma$$

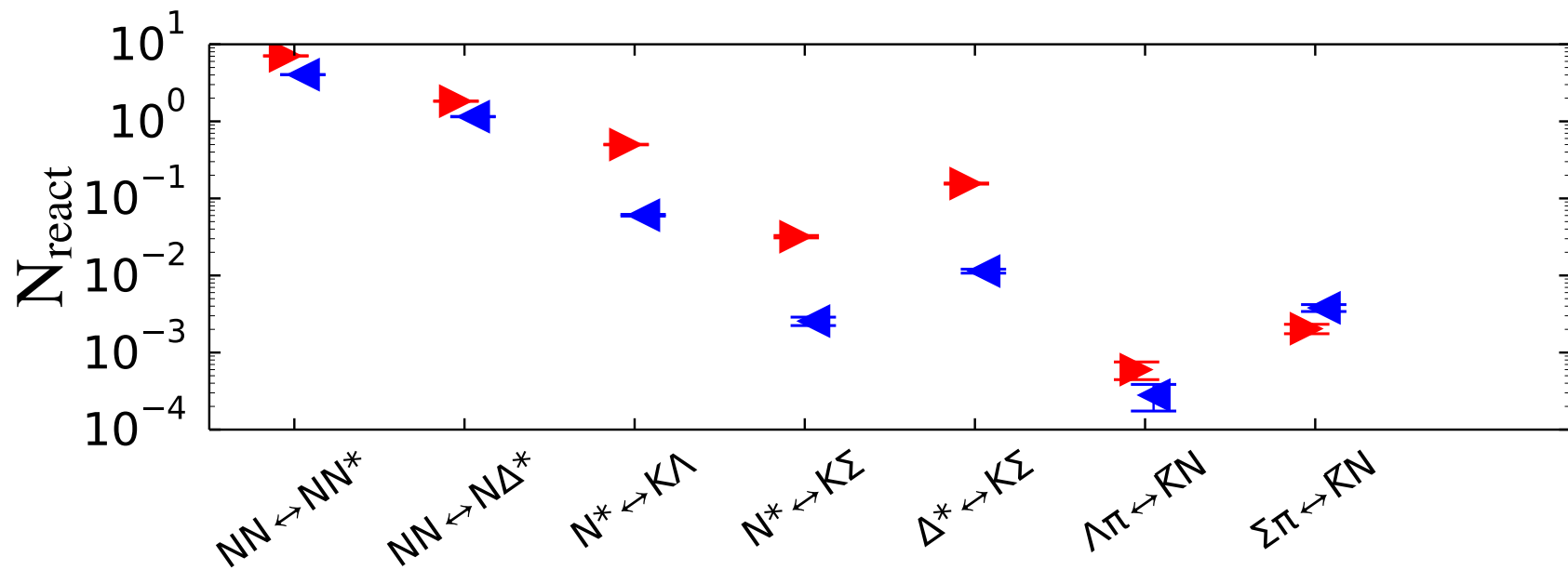
$$(3) \pi Y \leftrightarrow K^- N$$

(Gräf, Steinheimer et al. Phys.Rev. C90 (2014) 064909)

→ Only produced in secondary processes via resonances



Au +Au at $E_{\text{kin}} = 1.5 \text{ AGeV}$
 Events = 25000



Calculation specifics

- With Fermi motion
- No potentials
- No Pauli-blocking

- No string fragmentation necessary because $E_{\text{kin}} = 1\text{-}2 \text{ AGeV}$
- Results compared to measurements by
KaoS (Kaon Spectrometer)
HADES (High Acceptance Dielectron Spectrometer)

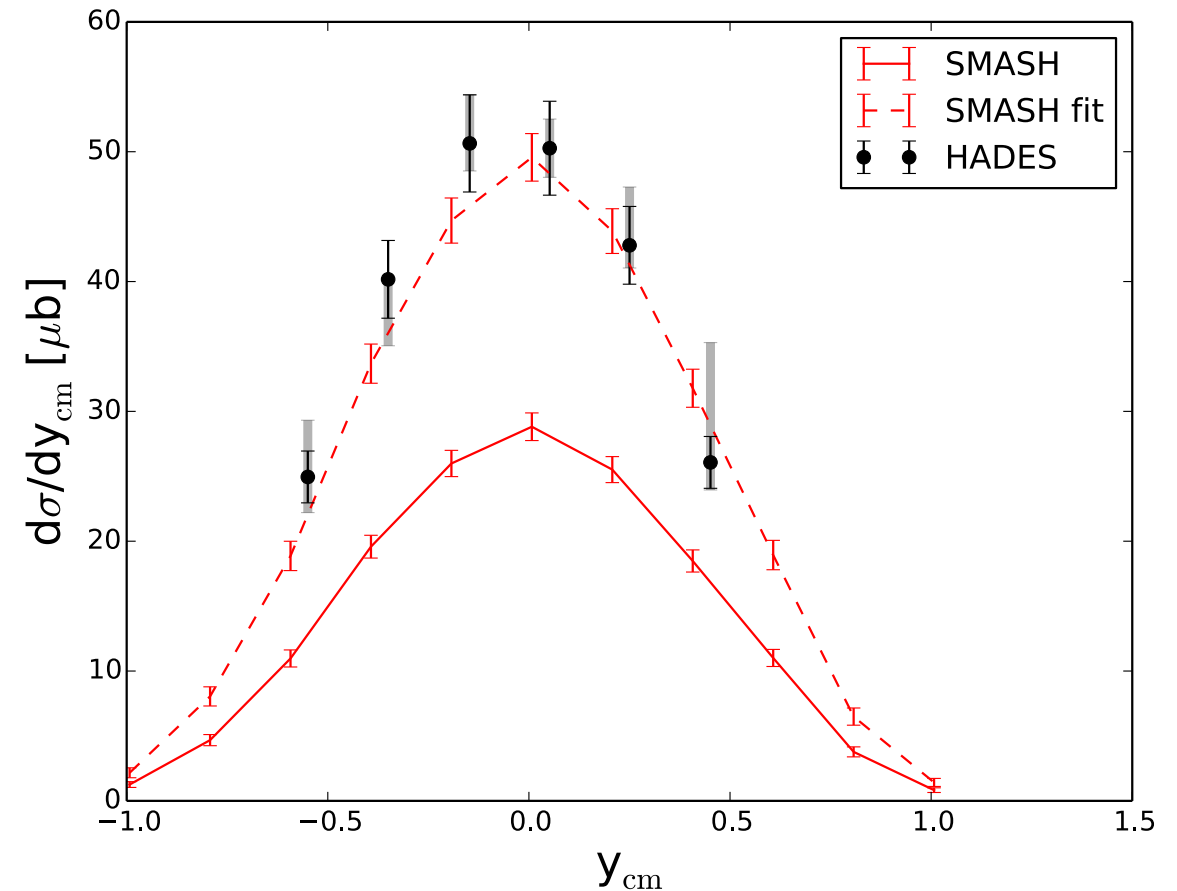
System	Observable
HADES: p+p at $E_{\text{kin}} = 3.5 \text{ GeV}$	Rapidity spectra y Transverse momentum spectra p_{T}
SMASH: Au+Au, $\pi^- + \text{C}$, $\pi^- + \text{W}$ at $E_{\text{kin}} = 1.7 \text{ AGeV}$	Multiplicities N
KaoS: Au+Au, Ni+Ni at $E_{\text{kin}} = 1.5 \text{ AGeV}$	Centrality dependence A_{part}
HADES: Ar+KCl at $E_{\text{kin}} = 1.76 \text{ AGeV}$	Transverse mass spectra m_{T}
KaoS: Au+Au at $E_{\text{kin}} = 1.5 \text{ AGeV}$	Effective temperature T_{eff}

p+p $E_{\text{kin}} = 3.5 \text{ AGeV}$, K_S^0

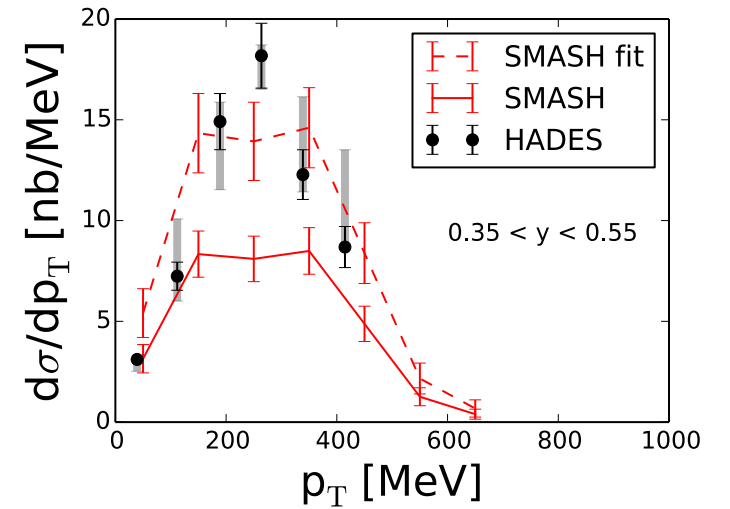
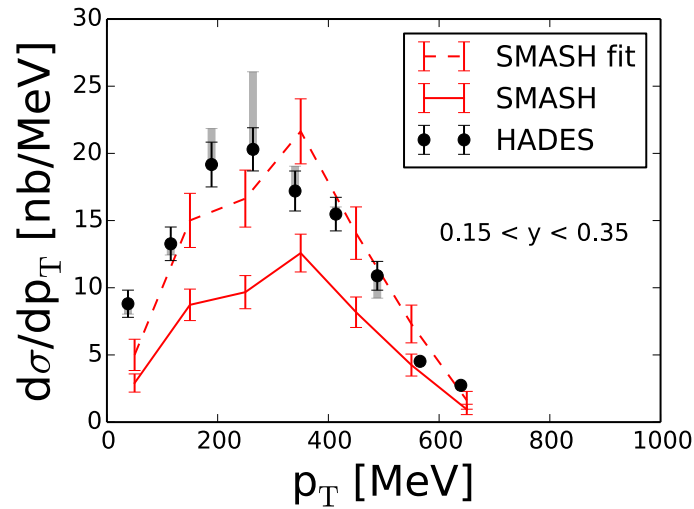
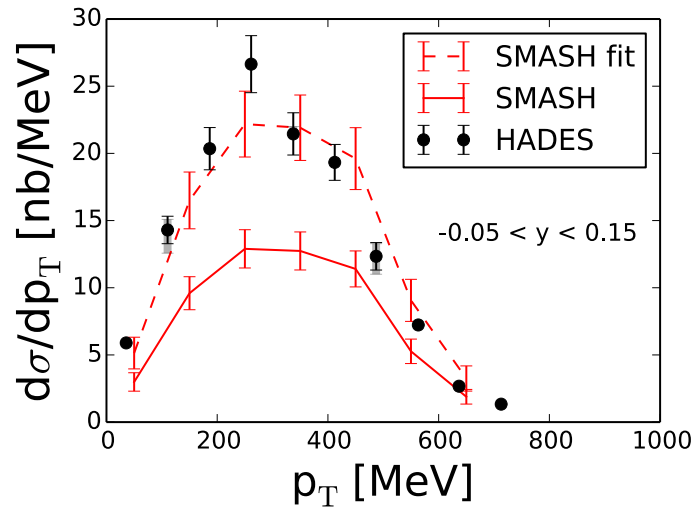
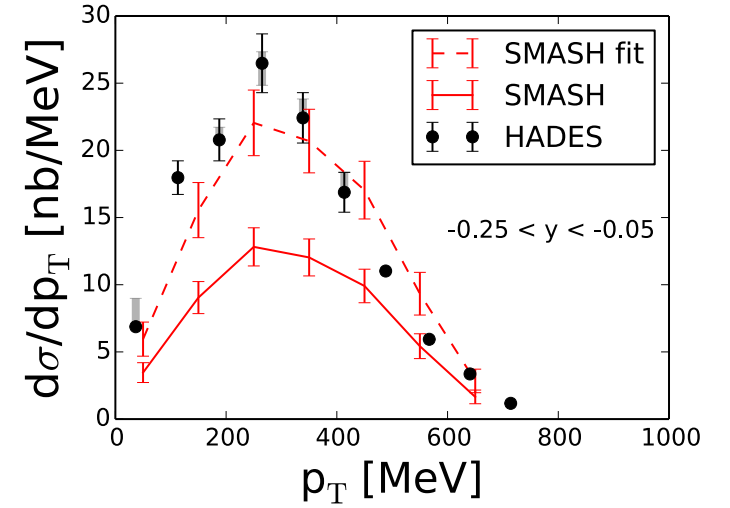
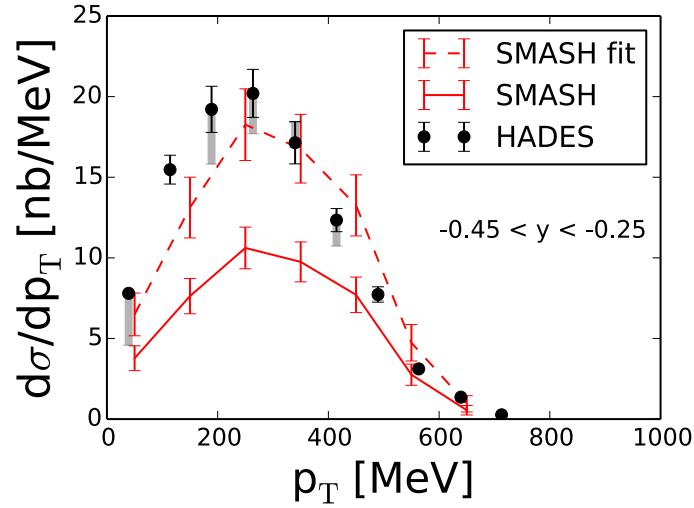
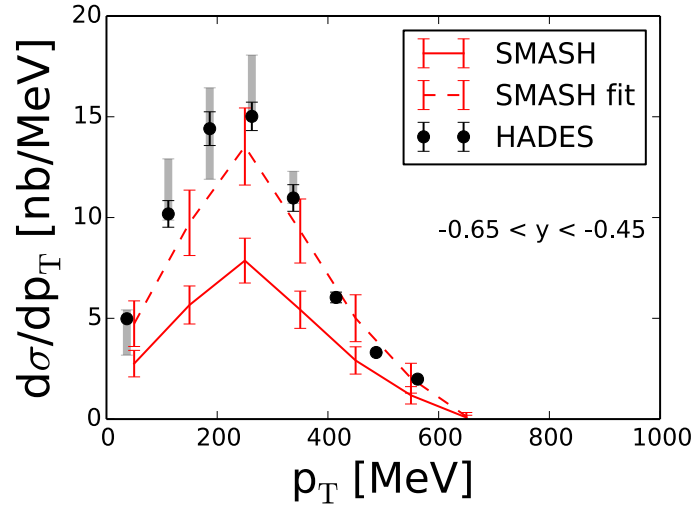
- K^0 and \bar{K}^0 form CP-eigenstates in K_S^0 and K_L^0
- Transverse momentum: $p_T = \sqrt{p_x^2 + p_y^2}$
- Rapidity: $y = 0.5 * \ln\left(\frac{E+p_z}{E-p_z}\right)$

p+p $E_{\text{kin}} = 3.5$ AGeV, rapidity spectra K_S^0

- Cross section calculated from properties of resonances
- $\sigma^{pp \rightarrow K^0 X} = 60 \mu\text{b}$ (SMASH)
- $\sigma^{pp \rightarrow K^0 X} = 103 \mu\text{b}$ (HADES)
(Agakishiev et al. Phys.Rev. C90 (2014) 054906)
- Dashed lines scaled by $\sigma^{pp \rightarrow K^0 X}$ ratio

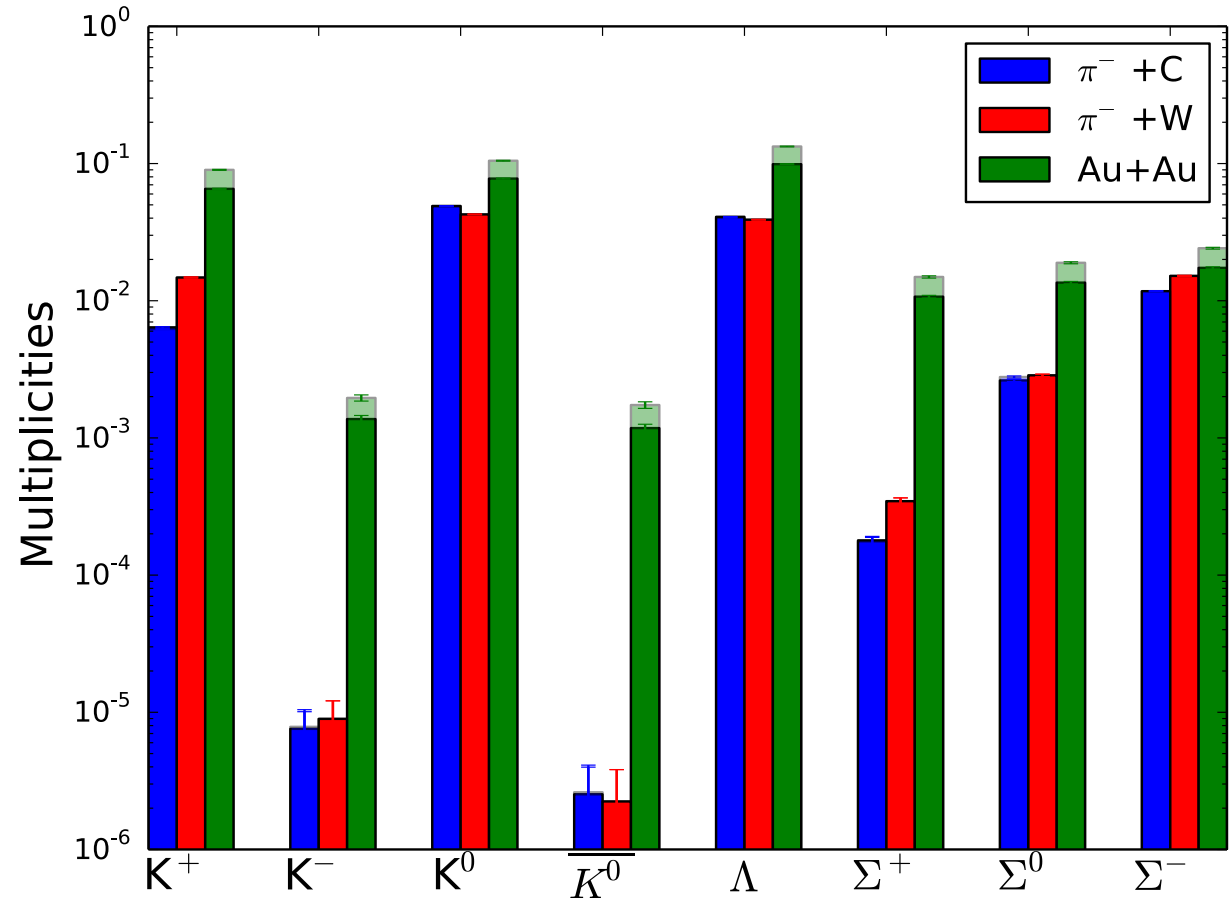


$p+p$ $E_{\text{kin}} = 3.5$ AGeV, p_T – spectra K_S^0



SMASH $E_{\text{kin}} = 1.7$ AGeV, Multiplicities

- Filled bars: without Fermi motion
- Transparent bars: with Fermi motion

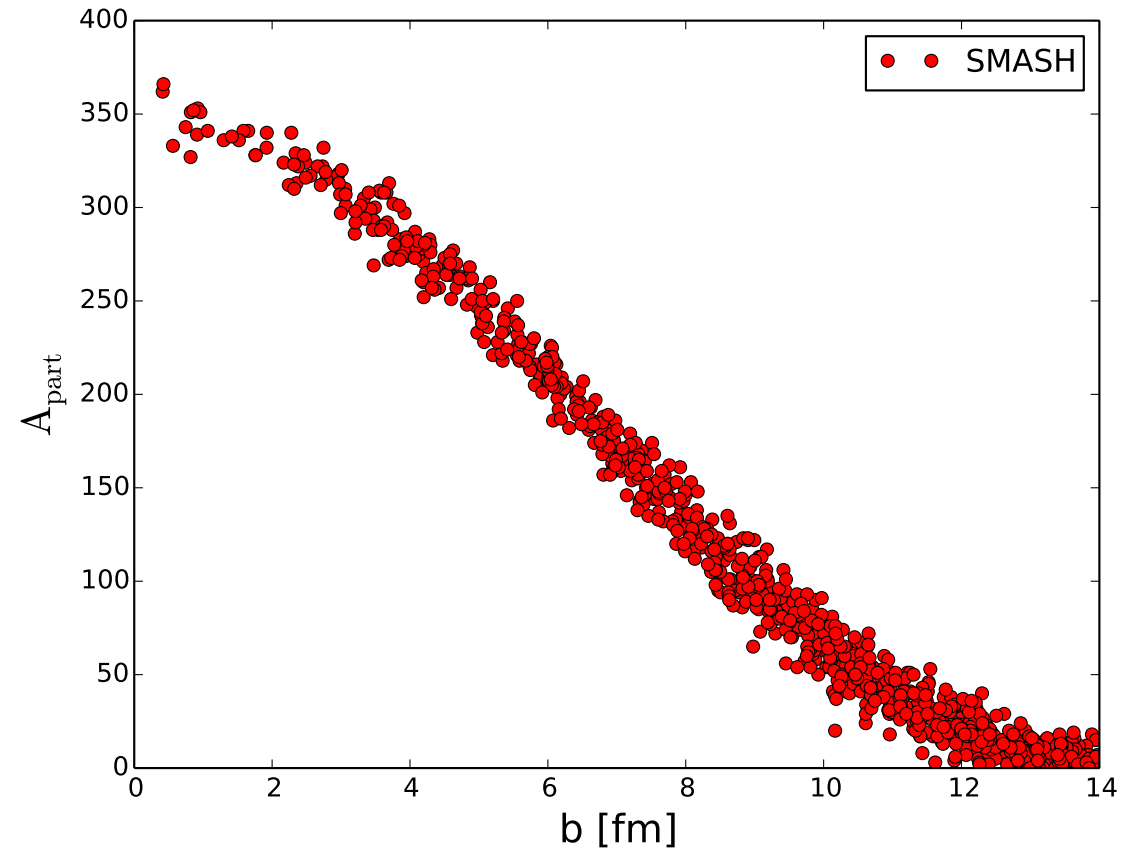


Heavy – ion reactions

Au+Au $E_{\text{kin}} = 1.5 \text{ AGeV}$

- Number of participants depending on impact parameter

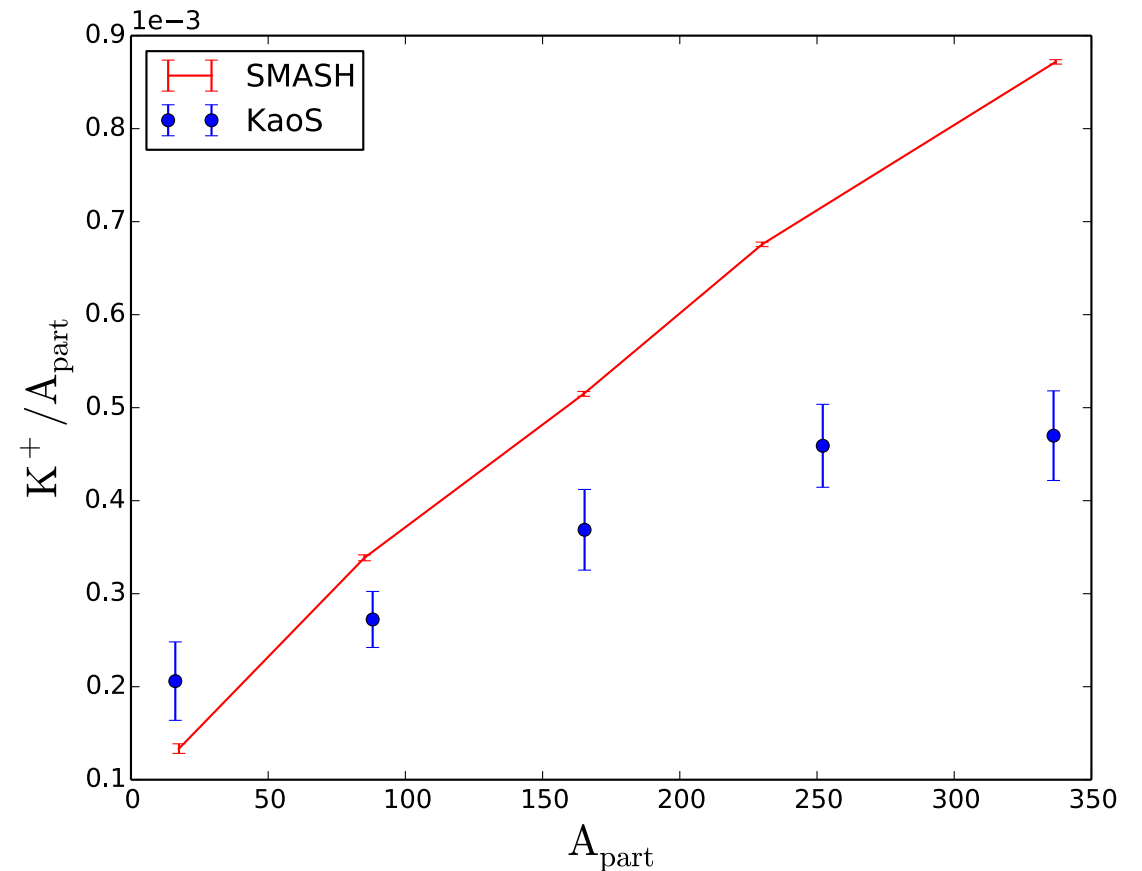
→ Number of participants decreases with higher impact parameter



Au+Au $E_{\text{kin}} = 1.5 \text{ AGeV}$

- Number of K^+ and K^- rises with A_{part}
- In comparison to KaoS : at high participants SMASH overshoots

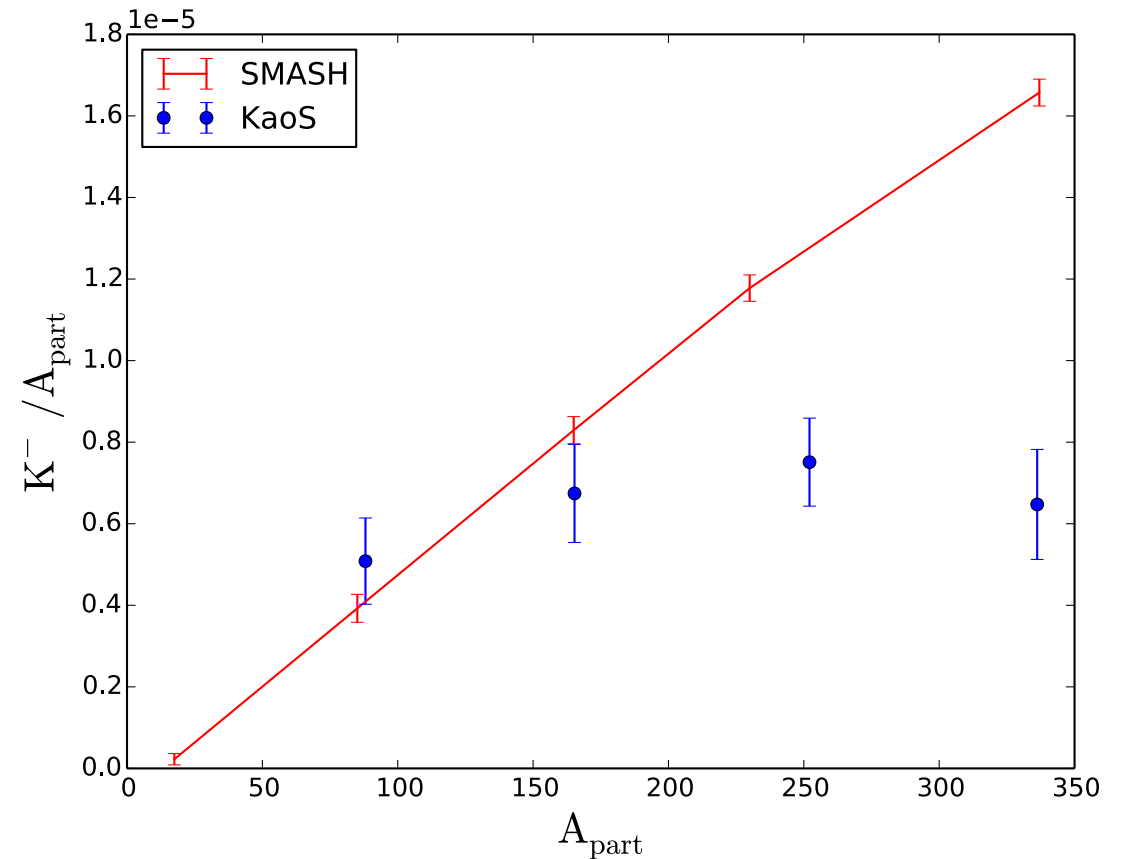
(A.Forster et al J.Phys. G31 (2005) no.6, S693-S700)



Au+Au $E_{\text{kin}} = 1.5$ AGeV

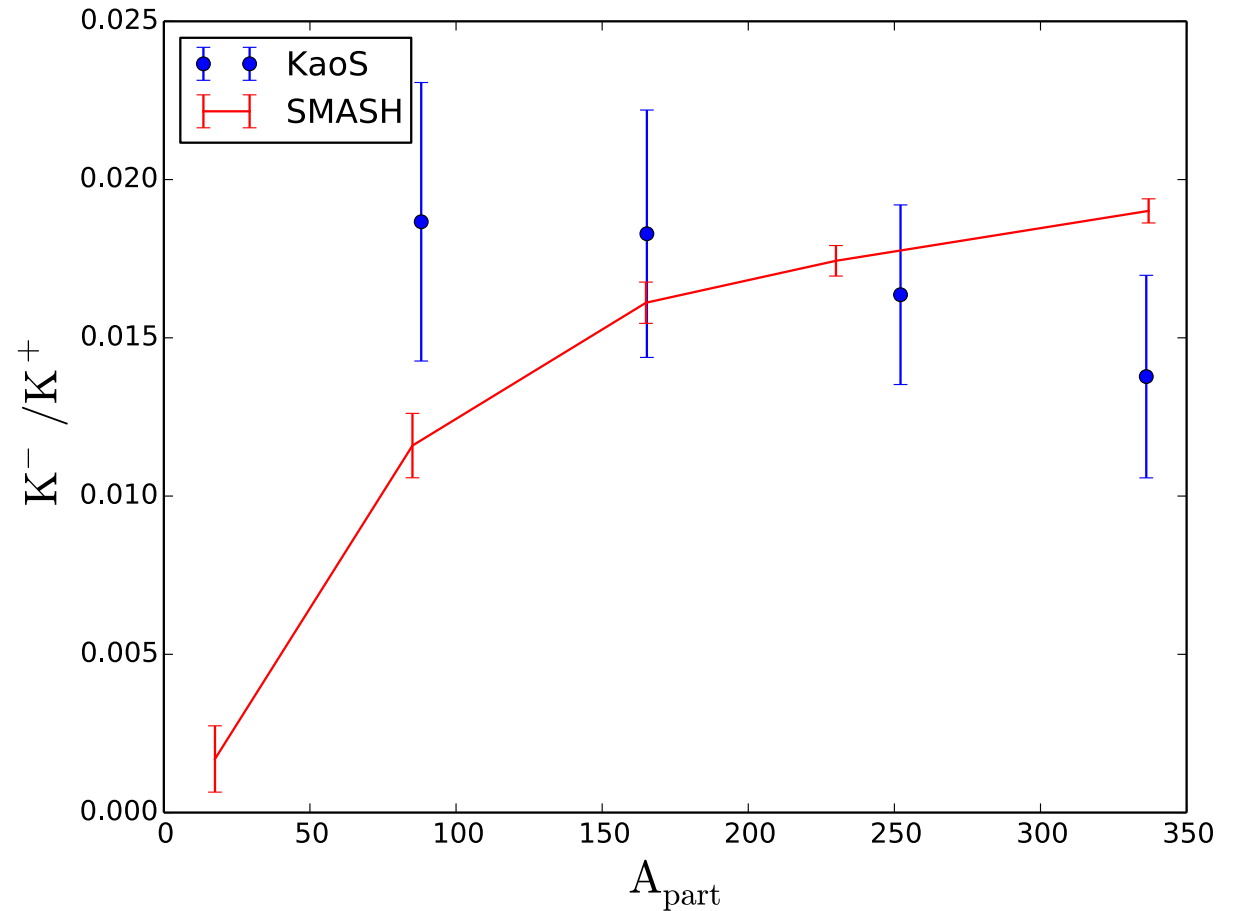
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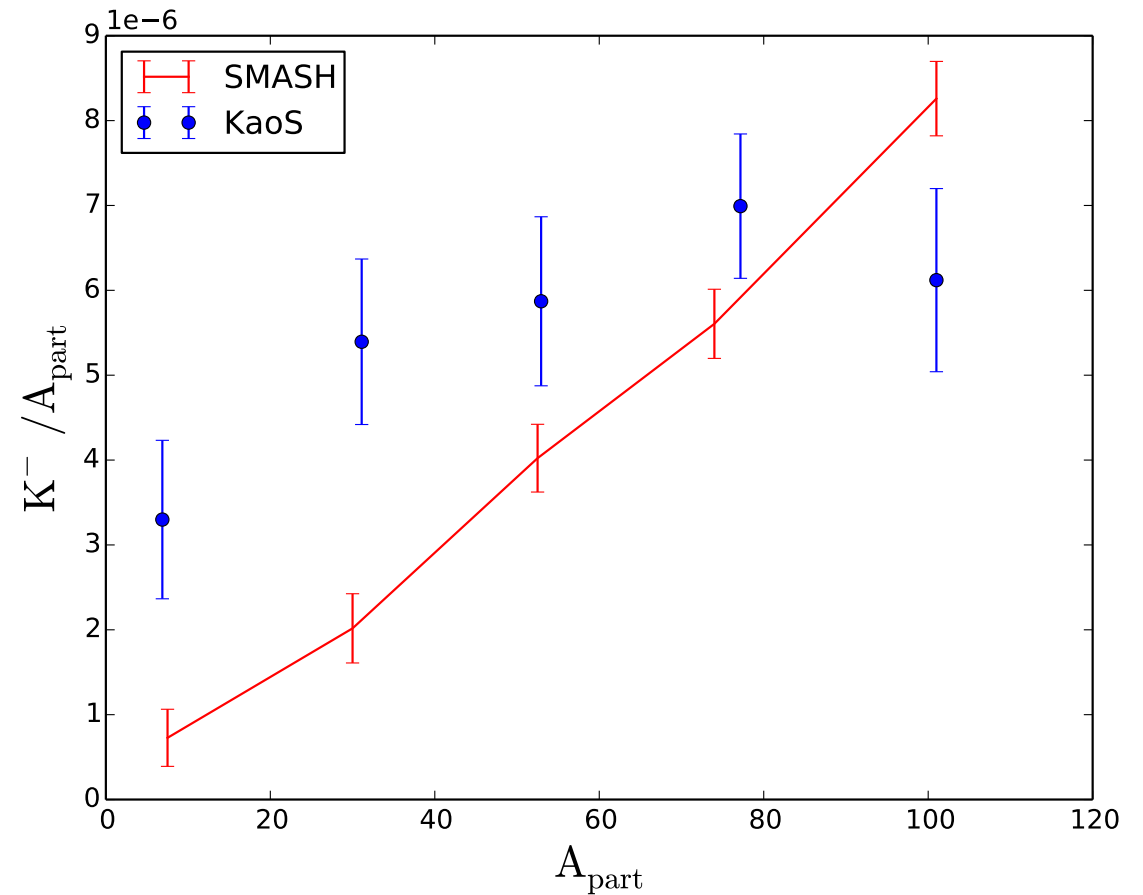
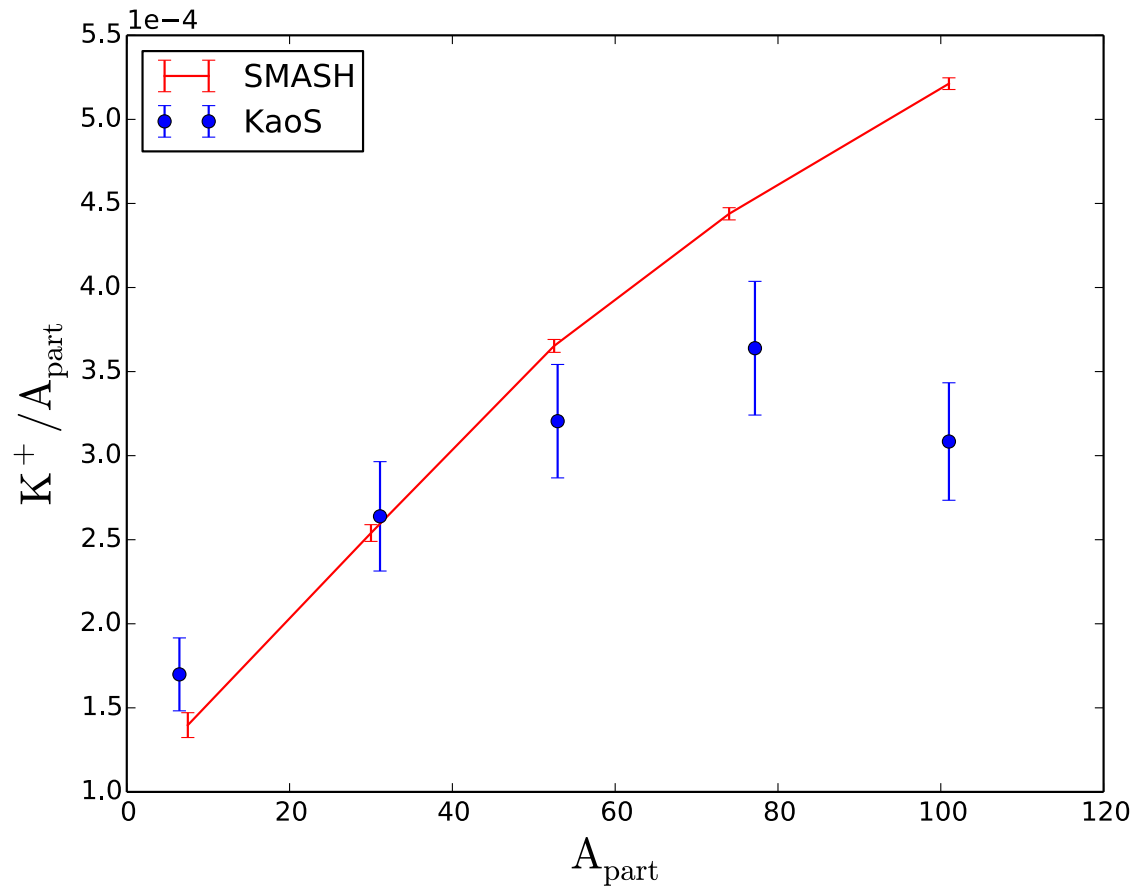


Au+Au $E_{\text{kin}} = 1.5$ AGeV

- KaoS
 - K^- to K^+ ratio is at 0.02
- SMASH
 - K^- to K^+ ratio increases with A_{part}



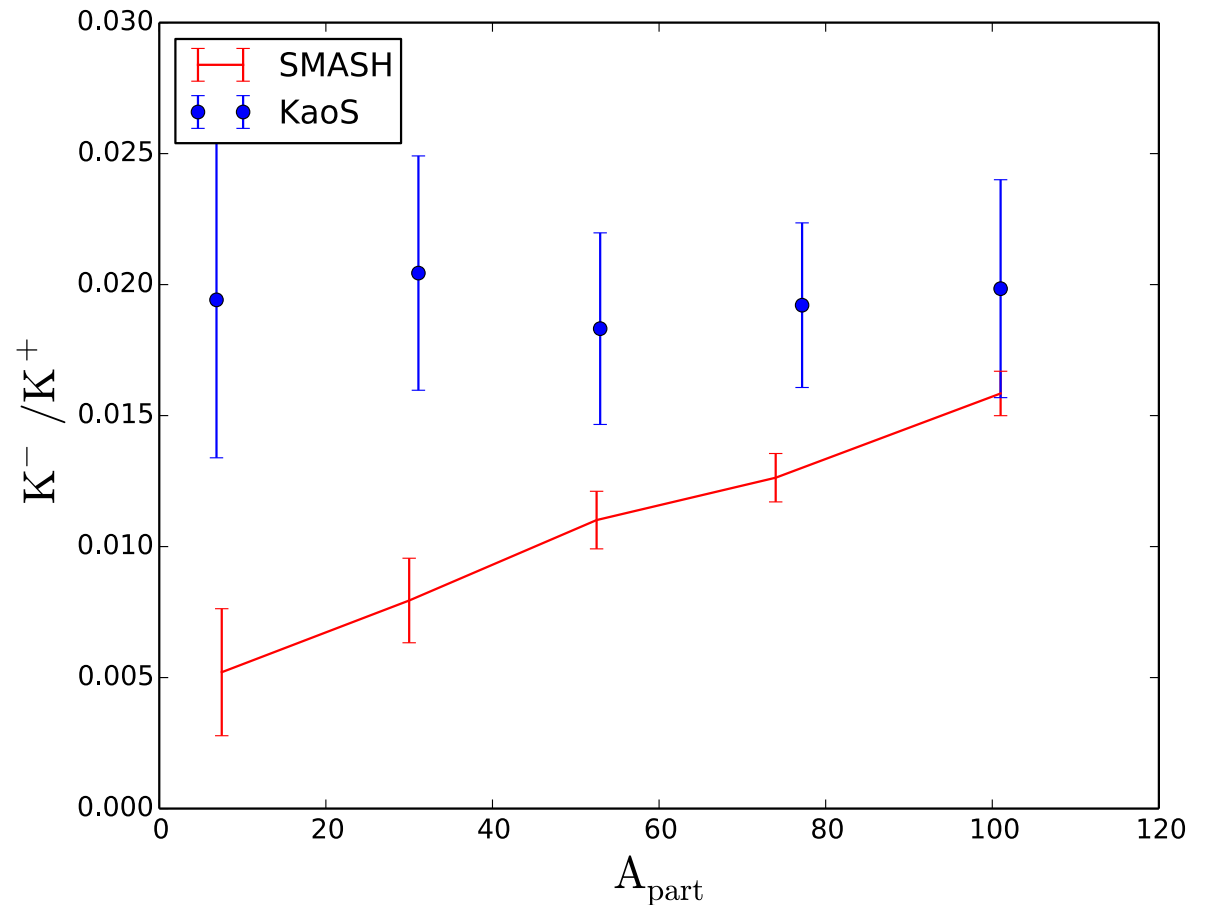
Ni+Ni $E_{\text{kin}} = 1.5 \text{ AGeV}$



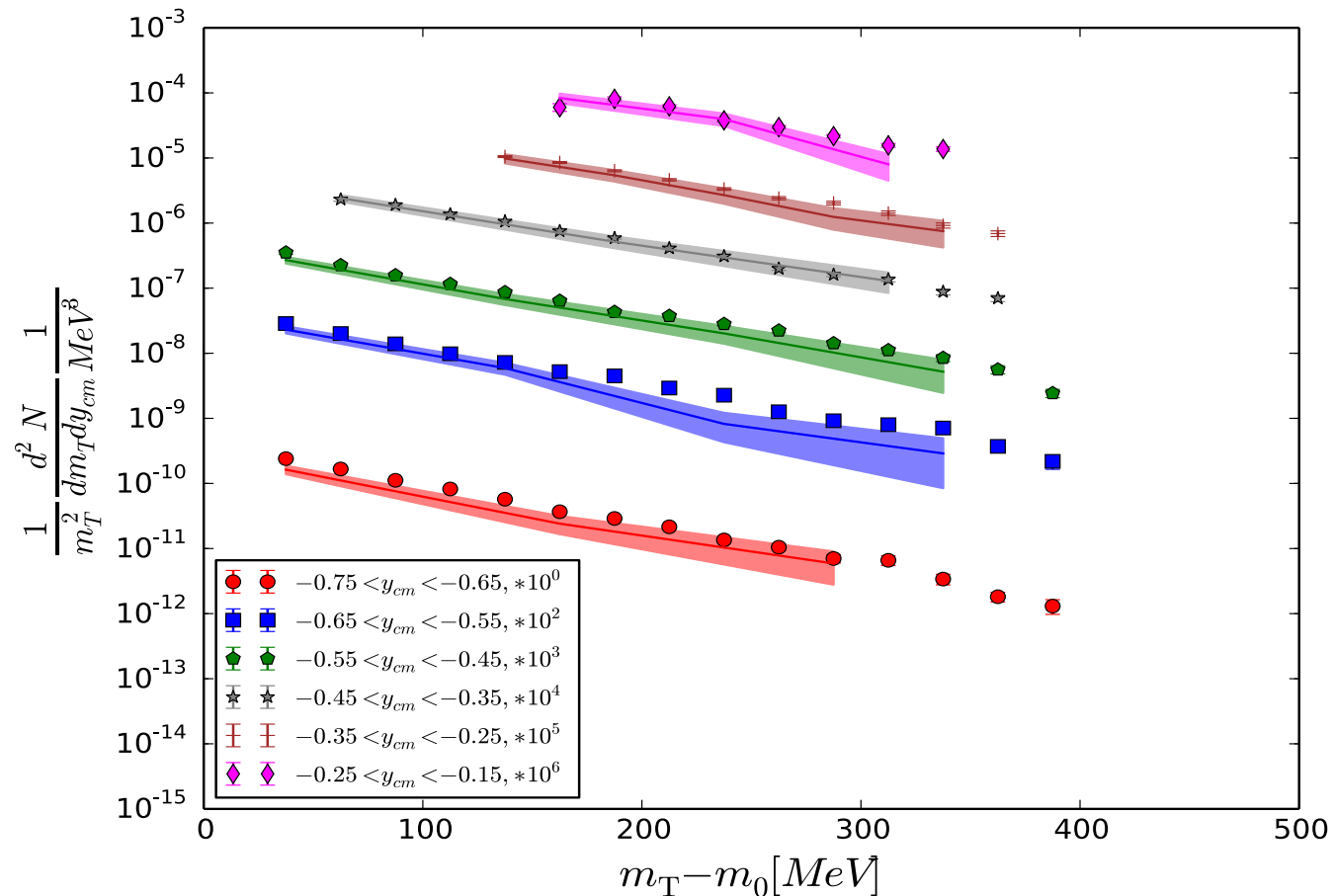
Ni+Ni $E_{\text{kin}} = 1.5 \text{ AGeV}$

- Number of K^+ , K^- and the ratio rises with A_{part}

→ Same behaviour for smaller systems



Ar+KCl $E_{\text{kin}} = 1.76$ AGeV, m_{T} – spectra K^+

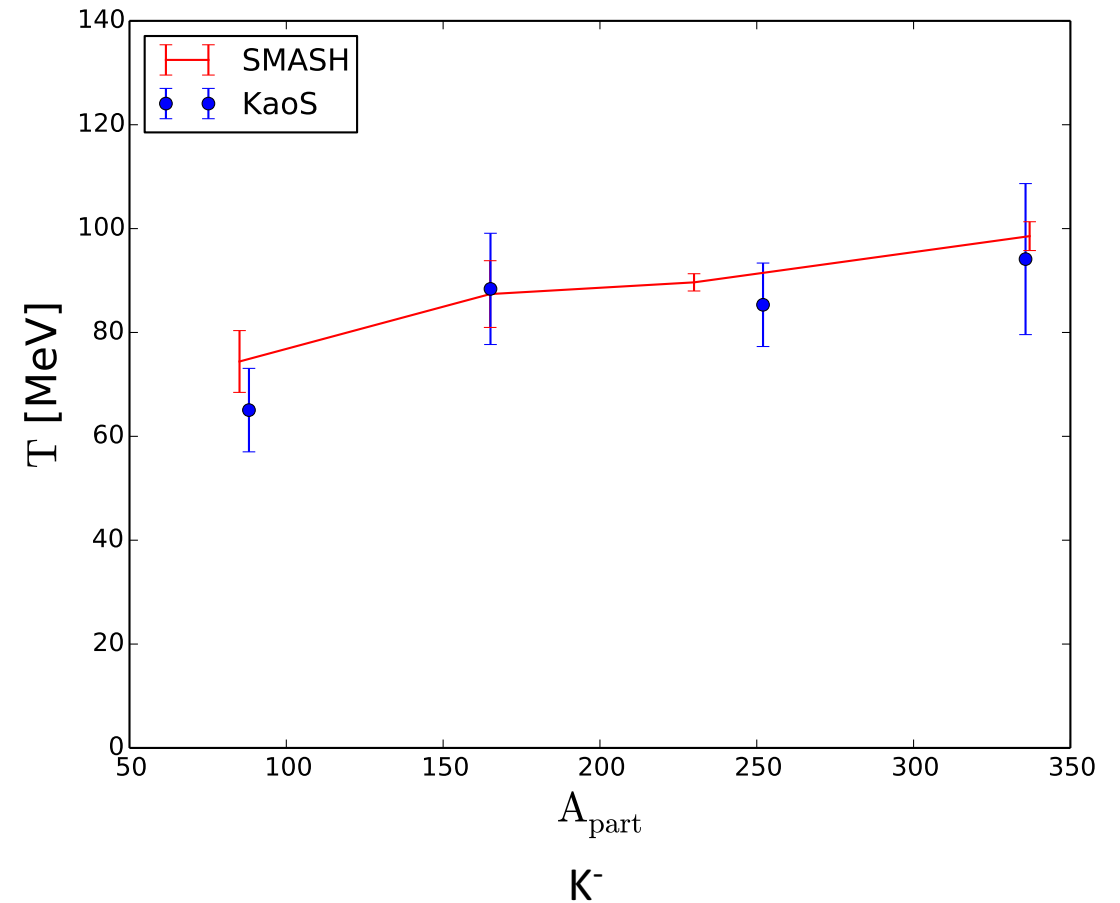
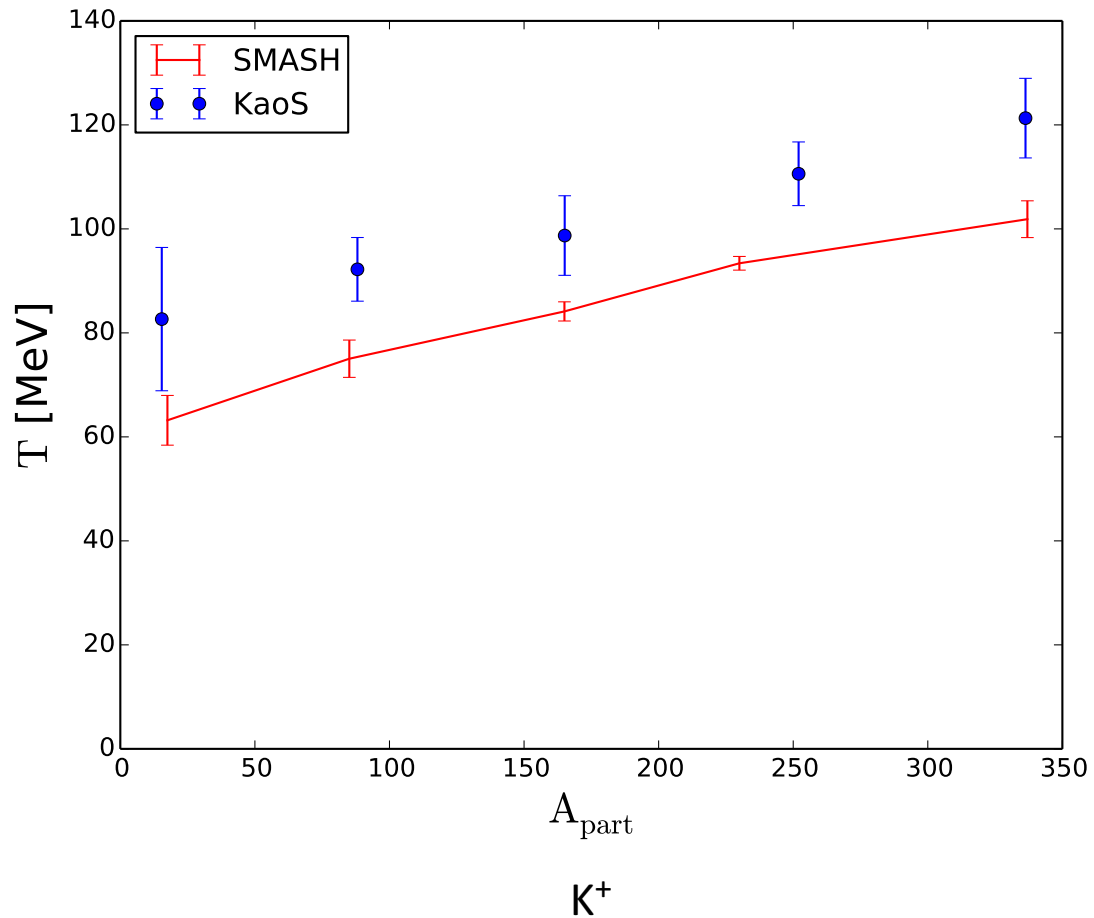


HADES = Ar + KCl
SMASH = Ca + Ca

Transverse mass distribution
has been determined in six rapidity bins,
ranging from:

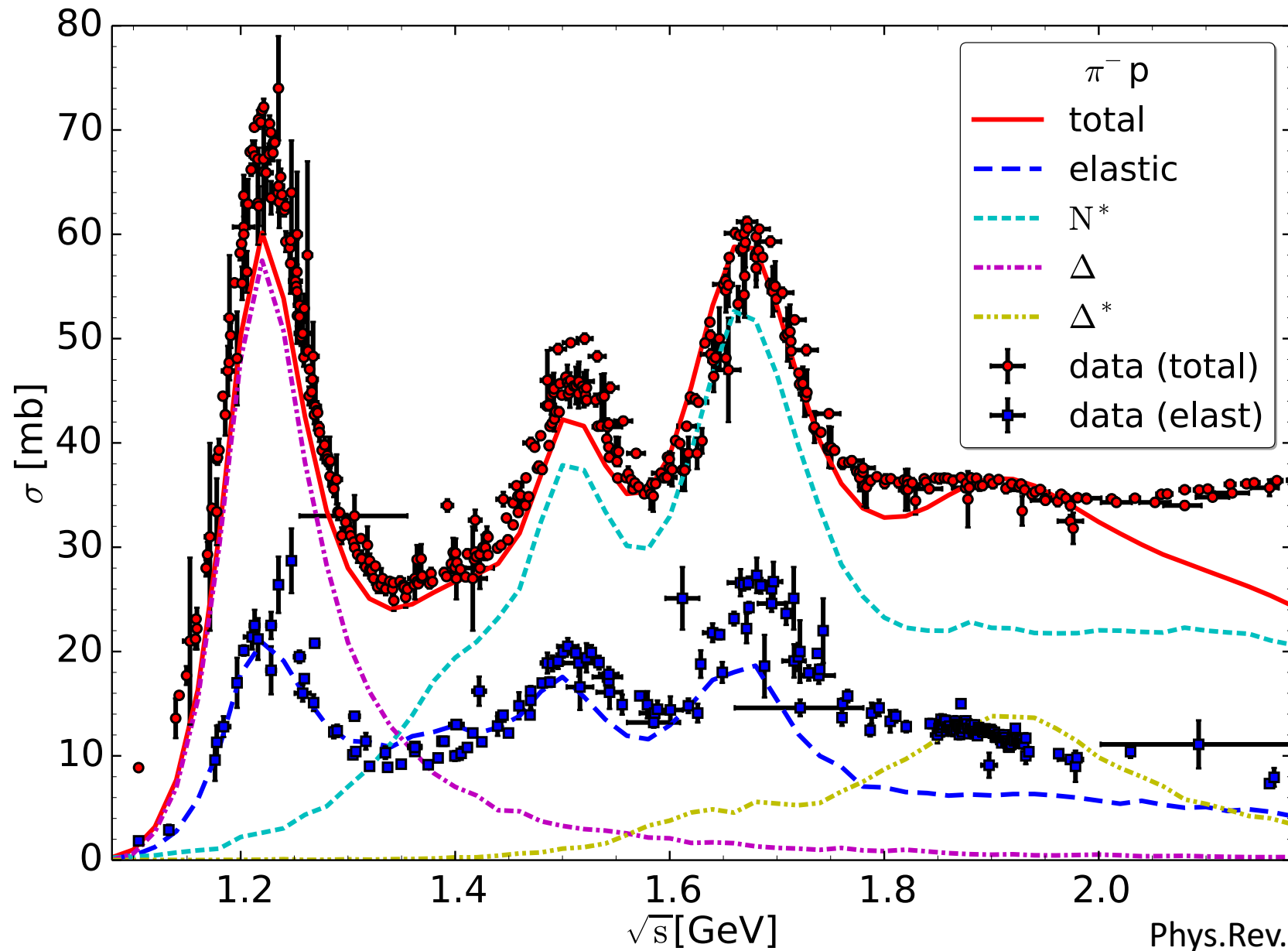
$-0.75 < y_{cm} < -0.15$

Au+Au $E_{\text{kin}} = 1.5 \text{ AGeV}$, T_{eff}

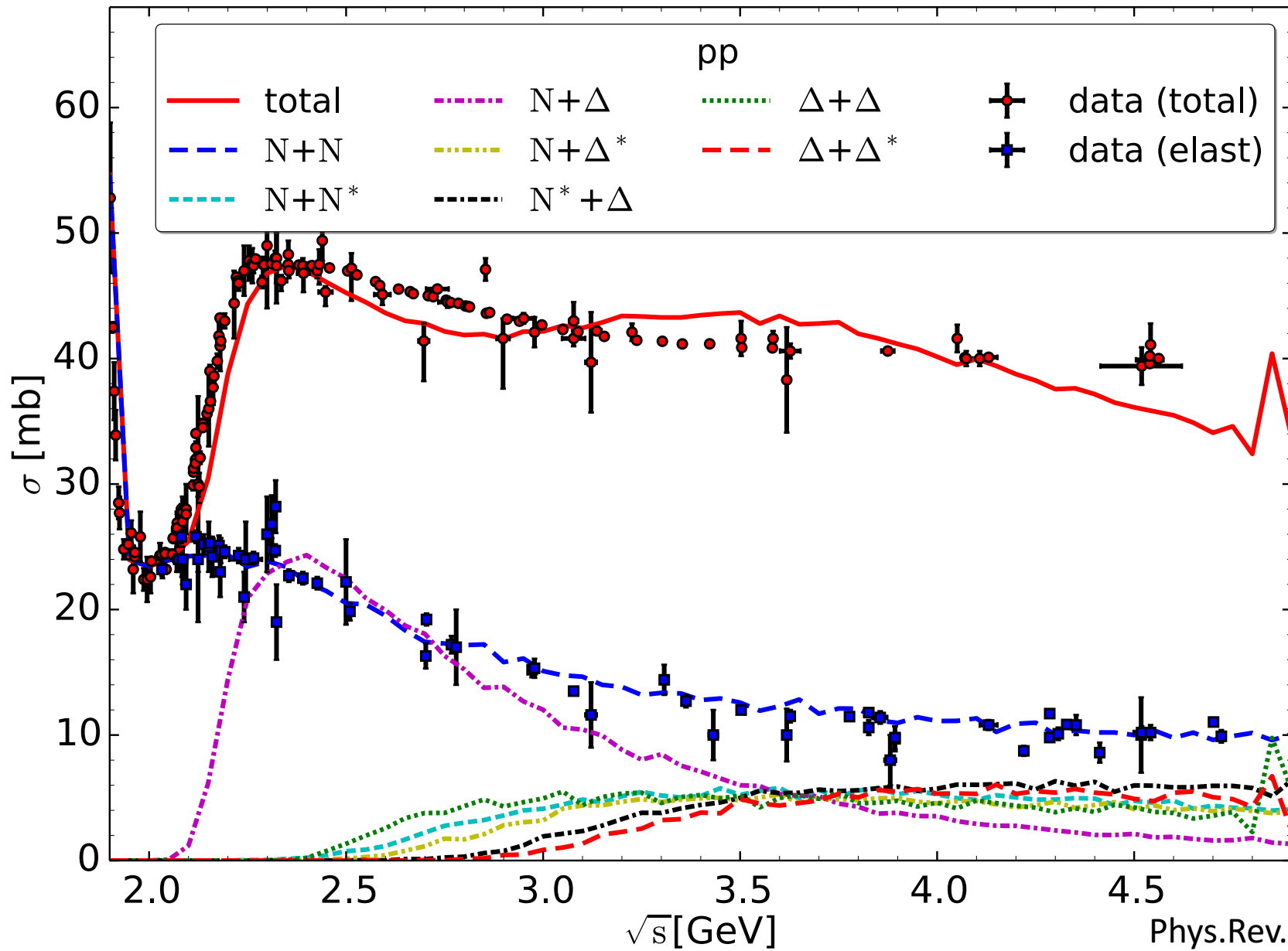


System	Observable	Results
HADES: p+p at $E_{\text{kin}} = 3.5 \text{ GeV}$	$\frac{d\sigma}{dy'}$, $\frac{d\sigma}{dp_T}$	K_S^0 : Lower multiplicities, same shape
SMASH: Au+Au, $\pi^- + C$, $\pi^- + W$ at $E_{\text{kin}} = 1.7 \text{ AGeV}$	N	Waiting for HADES data
KaoS: Au+Au, Ni+Ni at $E_{\text{kin}} = 1.5 \text{ AGeV}$	$\frac{K^+}{A_{\text{part}}}$, $\frac{K^-}{A_{\text{part}}}$, $\frac{K^-}{K^+}$	In central collisions: → Overshooting
HADES: Ar+KCl at $E_{\text{kin}} = 1.76 \text{ AGeV}$	m_T	Good agreement
KaoS: Au+Au at $E_{\text{kin}} = 1.5 \text{ AGeV}$	T_{eff}	K^+ : Lower K^- : Good agreement

Thank you for your attention!



Phys.Rev. C94 (2016) no.5, 054905



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